

Article

A Novel Approach to Generate Hourly Photovoltaic Power Scenarios

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Abstract: Photovoltaic power is playing an ever-increasing role in the energy mix of countries worldwide. It is a stochastic energy source, and simulation models are needed to establish reliable risk management. This paper presents a novel approach for simulating hourly solar irradiation and—as a consequence—photovoltaic power based on easily accessible data such as wind, temperature, and cloudiness. Solar simulations are generated via a multiplication factor that scales the maximum possible solar irradiation. Photovoltaic simulations are then derived using formulas that approximate the physical interdependencies. The resulting simulations are unbiased on an annual level and reasonably reflect historic irradiation movements. Interpreting our approach as a descriptive model, we find that error values vary over the year and with granularity. Errors are highest when considering hourly values in wintertime, especially in the morning or late afternoon.

Keywords: photovoltaic power; solar irradiation; simulation; cloudy sky model



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1. Introduction

Climate change is a major concern for the international community, with global warming being among the most visible effects [1]. In recent years, governments have been increasing efforts to fight climate change and global warming, and hence environmental degradation—challenges that confront modern society and cause numerous problems [2,3]. Thereby, the focus primarily lies on significantly reducing greenhouse gas emissions. Transforming energy production towards an (almost) carbon-neutral future becomes essential in this context. However, challenges are ample. Financial costs are considerable, and it remains challenging to integrate renewable energy into the electricity grid on a large scale [4,5]. The primary reason is that weather is stochastic, and so is wind or photovoltaic energy. Power utilities have to provide backups to prevent blackouts; companies and more minor participants such as individual households have to calculate with uncertain profits when deciding about investments in renewable energies such as rooftop solar panels.

Regarding photovoltaic power, there are numerous commercial and non-commercial tools for calculating the financial value of a solar panel. To the authors' best knowledge, these tools are exclusively based on historical weather observations and yield an average payoff for a panel. No risk analysis such as computing quantile values is possible for deciding about substantial financial investments with amortization periods of 10–20 years—an aspect that is becoming even more important in Germany as feed-in tariffs and, subsequently, profitability are decreasing.

The literature already offers some approaches to simulate photovoltaic power (see Section 2 and the discussion in Section 5.2), which is a challenging task. The output of a photovoltaic panel is influenced, among other things, by temperature, air pressure, aerosol

optical density, precipitation, ozone concentration, and humidity. Models cope with this complexity either by ignoring it or integrating it via adequate formulas. As a result, we have either univariate models limited solely to photovoltaic or irradiation data or fairly complex models requiring computational power and/or multiple input variables.

Univariate models do not explicitly model interdependencies between irradiation and variables such as temperature, which is required when using irradiation simulations in more complex scenarios such as optimizing a battery charging strategy. On the other hand, if too many input variables are required, we risk data problems regarding availability and/or quality data problems. This article now presents a novel model for risk management by providing Monte Carlo simulations for solar irradiation and relevant factors such as cloud coverage and temperature. Based on formulas from Myer's cloudy sky model (CSM) [6], we develop an understandable, straightforward, and convenient concept for deriving corresponding photovoltaic values and hence a basis for the risk analysis of a system whose payoff depends on solar irradiation. Our approach can be applied to any location worldwide.

There is no straightforward concept for evaluating the quality of our simulation model—especially regarding the photovoltaic simulations. Hence, we focus on irradiation values and the ability of our model to replicate true values. First, we checked both solar irradiation and temperature simulations for biases and found none, which is important as annual photovoltaic values can be considered reasonably accurate. Errors occur only regarding the distribution of solar hours over the year. Additionally, considering the shape of the simulated scenarios, these look reasonable compared to historical observations. When looking at the error levels, i.e., when interpreting our approach as a descriptive model, we see that error values vary with granularity and over the year. Relative errors of monthly irradiation amounts are maximally about 8%, whereas relative errors for hourly values vary between 17% and over 43%. We see the highest values in winter, especially in the margin hours (early morning, late afternoon). Note that these numbers have to be treated with care: Irradiation values in wintertime are comparably small, especially in the morning hours, so small absolute differences can result in larger relative errors. Eventually, errors are also higher for tilt solar panels (the common case) than for horizontal surfaces.

To sum up, when interpreting our approach as an explanatory model or forecasting model for solar irradiation, it is not an adequate alternative to existing approaches. However, our purpose is to simulate and not describe or forecast. Additionally, alternative models such as the Clear Sky Model of Bird and Hulstrom [7] require additional information such as aerosol density, which can be fairly hard to obtain, especially if historical data are needed. Moreover, we present reasonable weather simulations that are unbiased and are based on a model that can be implemented on almost any computer. It is ideal for off-grid systems in remote areas or for flexible off-grid systems such as the Origami photovoltaic panel suggested by Jasmin and Taheri [8]. An extensive comparison to other models can be found in Section 5.2.

The remainder of this paper is organized as follows. Section 2 presents a model for solar power production, while in Section 3 we discuss modeling temperatures and cloudiness. In Section 4 we analyze our model given real-world data. Section 5 contains a critical reflection of the results, whereby conclusions are drawn in Section 6.

2. Photovoltaic Power Production

For generating solar power scenarios, a stochastic model is needed. In literature, we see three alternative approaches:

- (A) **A direct model for photovoltaic power production.** One possibility is to model photovoltaic power as a univariate time series [9] or to apply a stochastic state-space model [10]. Despite the advantage of considering only a one-dimensional data set, this approach comes with a few challenges. We have multiple seasonality with constantly varying periods and amplitude (sunrise and sunset change every day) and sudden disruptions, e.g., due to rain (which might also have a seasonal impact).

- (B) **A separate model for all input parameters.** The second option is to first design separate models for all relevant influence factors, especially for irradiation and temperature. Given the respective numbers, such as the solar panels' tilt, we can then calculate the expected amount of photovoltaic power. Abdel-Nasser and Karrar [11] set up a model to forecast photovoltaic power based on neural networks and various input parameters. They ignore any knowledge about physical interactions and let the neural network decide. Miozzo et al. [12] apply a Markov process for simulation purposes. Barukčić et al. [13] propose an alternative stochastic approach. These ideas look promising because they are simple. Nonetheless, all models face the problem that temperature is relatively easy to model, but irradiation is not—mainly because of the dynamic seasonality described in Alternative (A).
- (C) **Physical deterministic models.** A physical deterministic model establishes a link between input factors such as temperature, cloudiness, albedo factor, and solar irradiation. Those factors are then modeled, including a stochastic component for each one, allowing simulations to be generated. Politaki and Alouf [14] proceed similarly by combining a deterministic model for solar power production under clear sky conditions with a Markov-based approach for cloudiness. A widely used fundamental model for modeling solar power is the clear sky model of Bird and Hulstrom [7], or the extension to cloudy conditions of Myers [6], which is accordingly called the Cloudy Sky Model (CSM). More complex solar irradiation models [15–17] for clear sky conditions are more accurate than Bird and Hulstrom's [7] model. However, they require a much larger number of input data, making them inadequate for daily use. Often, detailed live data for atmospheric input factors are missing or potentially skewed, which is the advantage of simplified models such as the one of Bird and Hulstrom [7]. The approach in Hofmann and Seckmeyer [18] also includes the possibilities of clouded skies and tries to balance complexity (it considers aerosol depth, for example) and simplicity. Additionally, they give a good overview of alternative approaches and benchmark them. For another summary of different concepts, please refer to Zhang et al. [5] or Khatib et al. [19], which focus on forecasting models.

The setup of the CSM is straightforward, easy to understand, and reasonable; only very few input factors are required. Moreover, those factors used in the CSM are much easier to model than solar radiation itself. Hence, we use the CSM as a basic model for our calculations.

2.1. Solar Power Production Modeling

Here, we introduce the relevant basics of solar power production. For an overview of how all the variables below are connected, please refer to Figure A1 in Appendix A. The output of electric power generated by a photovoltaic panel at time instance t , PV_t , is given by

$$PV_t = \Phi_t \cdot \eta_t \quad (1)$$

with η_t being the photovoltaic or solar panel's efficiency and Φ_t the solar flux. Thereby, η_t is time-dependent, as it is influenced by the panels' surface temperature $Temp_t^{surf}$. It can be modeled by a linear relationship between both variables [20,21]. In most cases, the relationship between η_t and $Temp_t^{surf}$ is derived from the respective photovoltaic panel data sheets [22,23]. For the modules produced by the German company *SolarEdge*, $\eta_t = 18.9 \cdot \left(100\% - 0.4\%/^{\circ}\text{C} \cdot \left(Temp_t^{surf} - 25^{\circ}\text{C}\right)\right)$. The surface temperature $Temp_t^{surf}$ is calculated via the total solar irradiation E_t^{tot} received by the photovoltaic system's surface of size A^S with an incidence angle Θ_t at hour t :

$$Temp_t^{surf} = Temp_t^{air} + 0.03 \left[\frac{^{\circ}\text{Cm}^2}{\text{Wh}} \right] \cdot E_t^{tot} \quad (2)$$

where $Temp_t^{air}$ denotes the average air temperature. The solar flux Φ_t is defined as

$$\Phi_t = E_t^{tot} \cdot A^S \cdot \cos \Theta_t. \quad (3)$$

The incidence angle Θ_t can be calculated (R package *solaR*) using information about the movements of earth and sun as follows [24]

$$\Theta_t = \arccos(\cos \Theta_t^Z \cdot \cos \beta + \sin \Theta_t^Z \cdot \sin \beta \cdot \cos(\gamma_t^S - \gamma)), \quad (4)$$

where β denotes the panel's tilt, Θ_t^Z the solar zenith angle at that specific time, γ_t^S the corresponding sun's azimuth angle, and γ the panel's azimuth. Note that Θ_t is also a zenith angle. The solar position algorithm of Michalsky [24] allows calculating all these values with a maximum error of $\pm 0.01^\circ$ in both solar zenith and azimuth angle [25], which is accurate enough for the application used in this work. This algorithm also calculates the extraterrestrial irradiation E_t^{extra} .

For horizontally arranged solar panels, E_t^{tot} equals the global irradiation at time t , E_t^{glo} , which is the sum of the direct irradiation E_t^{dir} and indirect or diffuse irradiation E_t^{dif} . Thereby, E_t^{dir} can be calculated based on E_t^{glo} [24], and $E_t^{dif} = E_t^{glo} - E_t^{dir}$ consequently. For tilted solar panels, we have to add reflected irradiation E_t^{refg} . Additionally, E_t^{dir} and E_t^{dif} are slightly changed to direct and diffuse irradiation on the tilted ground, noted by $E_t^{dir,tilt}$, $E_t^{dif,tilt}$ [26]. We can compute direct, diffuse, and reflected irradiation on the tilted ground using the Albedo $alb(\Theta_t^Z)$, which estimates how much irradiation is reflected from a surface (see Equation (6)). According to Alboteanu et al. [27], the formulas read

$$\begin{aligned} E_t^{dir,tilt} &= E_t^{dir} \cdot \frac{\cos \Theta_t}{\cos \Theta_t^Z}, \\ E_t^{dif,tilt} &= E_t^{dif} \cdot \frac{1}{2} \cdot (1 + \cos \beta), \\ E_t^{ref,tilt} &= E_t^{glo} \cdot \frac{alb(\Theta_t^Z)}{2} \cdot (1 - \cos \beta). \end{aligned} \quad (5)$$

The Albedo is exponentially increasing with a decreasing declination of the sun: the higher Θ_t^Z , the more light is reflected from the ground surrounding the panels. As it is difficult to obtain a functional form for the Albedo, we have to use approximations. We assume that the photovoltaic panels are mounted on roofs; hence, we assume an Albedo of 20% if the sun's declination is above a certain boundary [28]. Below this boundary, we approximate the Albedo using a linear function $alb(\Theta_t^Z)$:

$$alb(\Theta_t^Z) = \begin{cases} 20\% & \text{if } \Theta_t^Z \leq 60^\circ \\ 100\% - 8/3(90^\circ - \Theta_t^Z)\% & \text{else.} \end{cases} \quad (6)$$

2.2. How to Incorporate Clouds into the Model

The model of Bird and Hulstrom [7] (assuming clear sky conditions) allows computing solar irradiation on horizontal surfaces with an accuracy of $\pm 5\%$, while only basic atmospheric data are required. However, cloudy days are a lot more challenging to model. Myers [6] expands this model to the already mentioned CSM, where the relationship between irradiation under clear and under the cloudy sky is modeled via a multiplicative transmittance factor. Here, we propose a slight modification of this idea. We introduce two multiplicative factors, namely e^{glo} and $e^{glo,tilt}$, one for the global irradiation on the plain ground (to compute the Albedo influence) and one for the global irradiation tilted ground. The factors are used to establish a connection between extraterrestrial irradiation E_t^{extra} and irradiation on the ground under tilted conditions at time t :

$$\begin{aligned} E_t^{glo,clouds} &= E_t^{extra} \cdot e^{glo}, \\ E_t^{glo,tilt,clouds} &= E_t^{extra} \cdot e^{glo,tilt}. \end{aligned} \quad (7)$$

Using Equation (7), we can replace the first two formulas in Equation (5). The reflected irradiation under cloudy conditions, again, is computed using $E_t^{glo,clouds}$ from Equation (7) and the Albedo function (6) in Equation (5). Eventually, we yield E_t^{tot} under cloudy conditions as the sum of $E_t^{glo,tilt,clouds}$ and $E_t^{ref,tilt}$ under cloudy conditions. Via Equations (1)–(3), the power output PV_t is eventually computed.

3. Modeling Temperature and Cloudiness

Given the model design above, a stochastic model for temperature, cloudiness, and global (tilted) irradiation is needed to simulate solar power production scenarios. Temperature is a significant driver, as it influences the solar panels' efficiency. For this purpose, literature offers various so-called weather generators, which are models (mostly with corresponding software packages) that allow simulating different weather conditions [29]. As an alternative, a Fourier transform-based model is proposed [30]. More advanced (multivariate) approaches have one thing in common [31–33]: They contain a term that explicitly models the seasonal component. Most are autoregressive approaches, which is reasonable as the temperature will not change too much from one hour to the next. In most cases, a Gaussian distributed noise term is applied to include stochastic elements. The advantage of the autoregressive models is evident: They are easy to understand and handle, which is beneficial in practical application. Hence, we apply such a model to describe and simulate air temperature $Temp_t^{air}$. A sine function captures the trend. Using the Fourier transform delivers slightly better results but adds complexity to the model. Moreover, it is usual to compute the daily average temperature using a sine or cosine-based formula in business. Accordingly, let S_t be the annual seasonality on an hourly basis defined as

$$S_t = a + b \cdot \sin\left(\frac{(t - c)2\pi}{365.25 \cdot 24}\right), \quad (8)$$

where t denotes the hourly time index and $a, b, c \in \mathbb{R}$ are parameters to shift and scale the seasonal trend. To consider leap years, we write 365.25 instead of 365. The discrete-time hourly air temperature model itself is defined as

$$\begin{aligned} \Delta Temp_t^{air} = & \kappa_1(S_t - Temp_t^{air})\Delta t + \kappa_2(S_{t-1} - Temp_{t-1}^{air})\Delta t \\ & + \kappa_3(S_{t-24} - Temp_{t-24}^{air})\Delta t + \sigma\epsilon_t, \end{aligned} \quad (9)$$

where $\kappa_1, \kappa_2, \kappa_3 \in \mathbb{R}$, $\epsilon_t \sim N(0, 1)$, and $\sigma > 0$ denotes the volatility. If desired, one could design volatility to be time-varying, i.e., σ is replaced by σ_t (for potential models, refer to McNeil et al. [34]). The model itself is constructed to be mean-reverting around the annual seasonality S_t , which is reasonable, as there might be warm and cold winters, but the average temperature in December is lower than in July. Additionally, we have an autoregression lag of 1 and 24. Lag 1 accounts that temperature usually is not significantly changing from one hour to the other, and Lag 24 is the daily seasonality. This seasonality is sometimes more robust and sometimes weaker (in case of temperature drops, for example), but it is there. On average, it is warmer at noon than at midnight. The larger κ_3 , the larger the effect of daily seasonality.

Cloudiness, again, is measured in Germany as discrete states between 0 (clear sky) and 8 (a dense blanket of clouds). In other countries, it is recorded in the percentage of sky covered. Cloud movement is very complicated to model, and as it is not the paper's focus, we use a rather hands-on, data-driven approach. As distribution for the cloudiness, we propose using a conditional multinomial distribution conditional on the previous state to account for autoregressive effects. The transition probabilities are estimated based on a data set of measured cloudiness. Let $c(t) \in C$ be the cloudiness at a specific hour and let $C = \{c_1, c_2, \dots, c_k\}$, $k \in \mathbb{N}$ denotes the set of possible states of cloudiness. The conditional

probability $p_{i,j} = P(c(t) = c_i | c(t-1) = c_j)$ of having a cloudiness level of $c(t) = c_i$ given that the cloudiness of the previous hour was $c(t-1) = c_j$ is estimated as follows

$$p_{i,j} = \frac{\#days \text{ with cloudiness } c_i \text{ given } c_j \text{ on the previous day}}{\#days \text{ where previous day had a cloudiness of } c_j}. \quad (10)$$

The result is a matrix of transition probabilities, making it very easy to simulate future cloud coverages.

The transmittance factors $\epsilon^{glo,clouds}$ and $\epsilon^{glo,tilt,clouds}$ from Equation (7) are modeled as linear functions depending on T_t^{air} , $c(t)$, incident angle Θ_t^Z , the current hour, and E^{extra} . A seasonality parameter is also included. For each hour, we fit the model

$$y_t = \alpha_0 + \alpha_1 Temp_t^{air} + \alpha_2 c(t) + \alpha_3 E_t^{extra} + \alpha_4 (90^\circ - \Theta_t^Z) + season_t + \gamma_t \quad (11)$$

using least squares estimation. Thereby, $\alpha_0, \alpha_1, \alpha_2, \alpha_3, \alpha_4 \in \mathbb{R}$ represent the linear coefficients, $\gamma_t \sim \mathcal{N}(0, \sigma^2)$ the residuals, and y_t is either $\epsilon_t^{glo,clouds}$ or $\epsilon_t^{glo,tilt,clouds}$. Moreover, $season_t = -\cos(month(t) \cdot 2 \cdot \pi / 12)$, whereby $month(t)$ denotes the current month in numbers, i.e., March = 3. Note that we only obtain estimates for positive values of $(90^\circ - \Theta_t^Z)$, i.e., when the sun is above the horizon. For the calibration, historical values are used to obtain $E_t^{glo,tilt,clouds}$ via (7). As mentioned above, E_t^{extra} can be computed according to Michalsky [24]. Mind that the parameter estimates of Equation (11) vary with β and the azimuth choice. Hence, for each panel position, a new calibration is required.

4. Application to Real-World Solar Irradiation Data

For a summary of the simulation procedure, please refer to the flow chart in Figure A2 in Appendix A. Here—because it is the novel aspect of our model, and in order to keep the article short—we solely discuss the irradiation model. Given proper irradiation simulations, corresponding photovoltaic simulations can be derived as described in the flow chart (Figure A2).

4.1. Data

We use the Climate Data Center provided by the German Climate Service www.dwd.de (accessed on 1 February 2021) [35] for retrieving temperature and cloud coverage data. Irradiation data come from the CAMS Radiation Service v3.2 all-sky irradiation provided by MINES ParisTech, Paris, France. The acronym CAMS stands for Copernicus Atmosphere Monitoring Service, which provides data at a resolution of about 3 km–max. 5 km [36]. All data were accessed via www.soda-pro.com on 1 February 2021 [37]. An overview of the data sets, including length and sources, is given in Table 1.

Table 1. Data sets (all accessed on 1 February 2021).

Data Set	Source	From/To	Length
Temperature Ulm	www.dwd.de [35]	1 January 2014–31 December 2018	37,924
Cloud Cover Ulm	www.dwd.de [35]	1 January 2014–31 December 2018	37,381
Irradiation Ulm	www.soda-pro.com [37]	1 January 2013–31 December 2018	52,584

In Figure 1, we exemplarily plot one year (2018) of global solar radiation. We can see the expected seasonal swing and some substantial variation, making prediction quite hard and requiring the development of various scenarios to assess a financial investment influenced by irradiation values.

The daily pattern can only be conjectured; hence, we plot the patterns in two sample months in Figure 2. Now, the daily structure and the weather-related variation can be observed. There is a distinct daily seasonality; its extent is not constant and influenced by various weather phenomena. Figure 3, again, presents temperature and cloud cover data

for 2018. The stochastic pattern of the cloud cover is observed. Temperature shows the expected characteristic annual swing with a significant drop in the first quarter of 2018 and some significant variation over the year. There is also some daily seasonality, which is not constant due to changing weather and events such as sudden temperature drops. A stochastic model needs to incorporate all those observations.

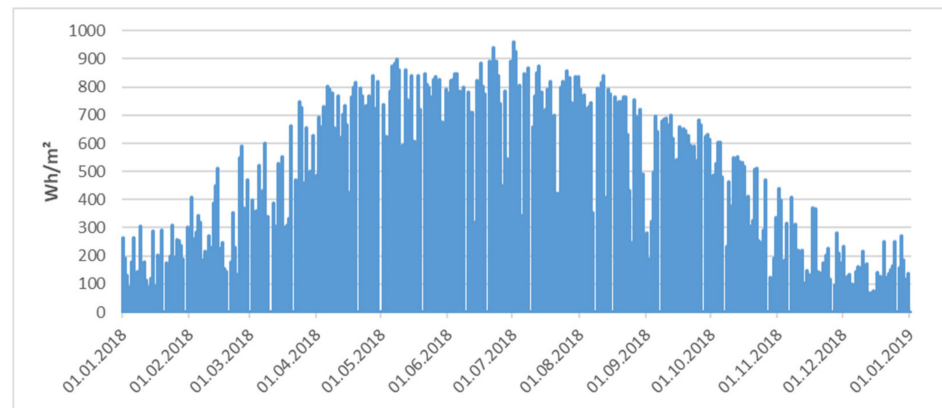


Figure 1. Hourly global solar irradiation from Ulm 2018.

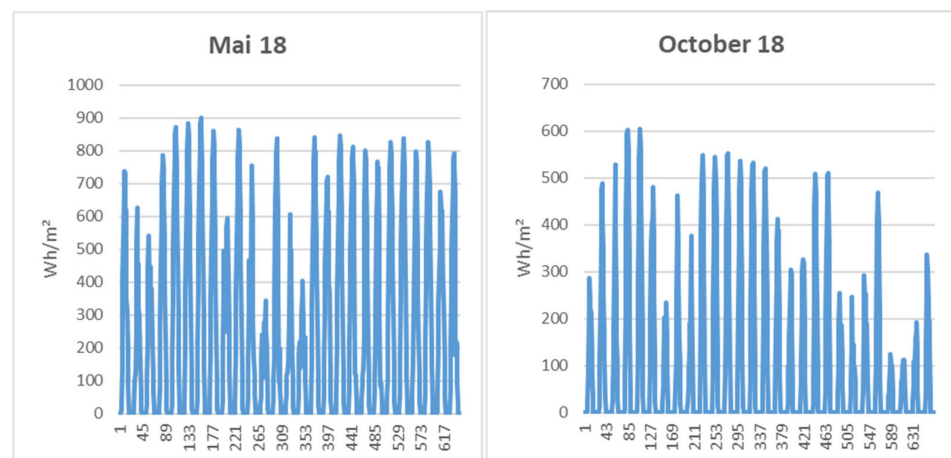


Figure 2. Hourly global solar irradiation from Ulm in May and October 2018.

4.2. Model Calibration and Results

We calibrate the model from Equation (9) to the above data using nonlinear least-squares estimation to generate temperature simulations. This model includes both an annual and a daily seasonality and allows for enough variation to simulate temperature drops or unusually low or high temperatures. Again, transition probabilities for the cloud coverage are computed using the available data sets according to Equation (10). For detailed results, please refer to Table A1 in Appendix A. The transmittance coefficients are computed using Equation (11) calibrated to the given data sets. Parameter estimates are given in Tables A2 and A3 in Appendix A.

In Figure 4, we exemplarily plot modeled hourly irradiation data for 2018 compared to the true values. It is not a simulated set but a descriptive model's output. It shows that the model captures the annual swing, and values are in the range of the historical observations. On a first glimpse, our model looks a bit spikier. However, this might also be due to the resolution. In Figure 5, we show the results for February (as a winter month) and July (as a summer month) in detail. Note that values are computed for photovoltaic panels with a 30° southward tilt. We can see that the relevant dynamics are captured in both months, whereby our model produces a smoother version, which is due to the least-squares fitting

of the model parameters in Equation (11). This is distinctly visible in the February data, where the model generates a smoothed version of reality.

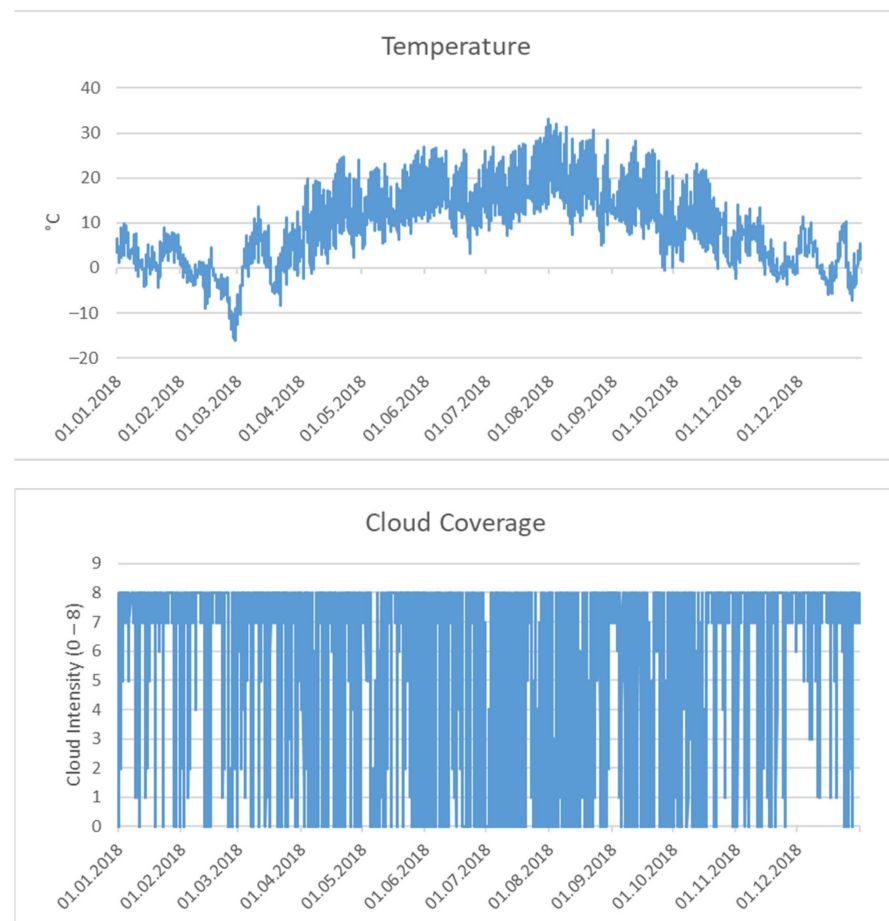


Figure 3. Hourly temperature and cloud cover data from Ulm in 2018.

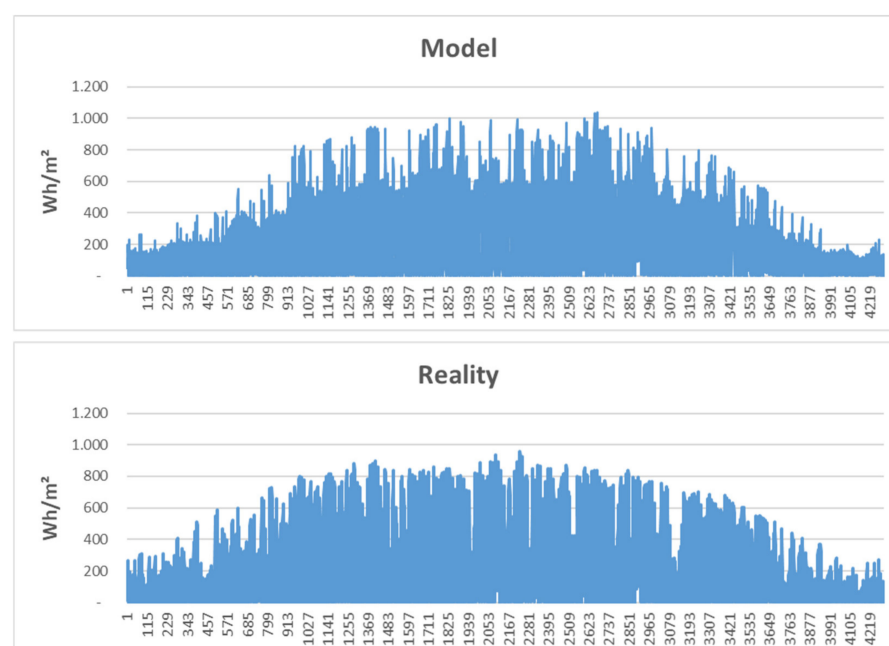


Figure 4. Hourly irradiation data for 2018: our model vs. reality.

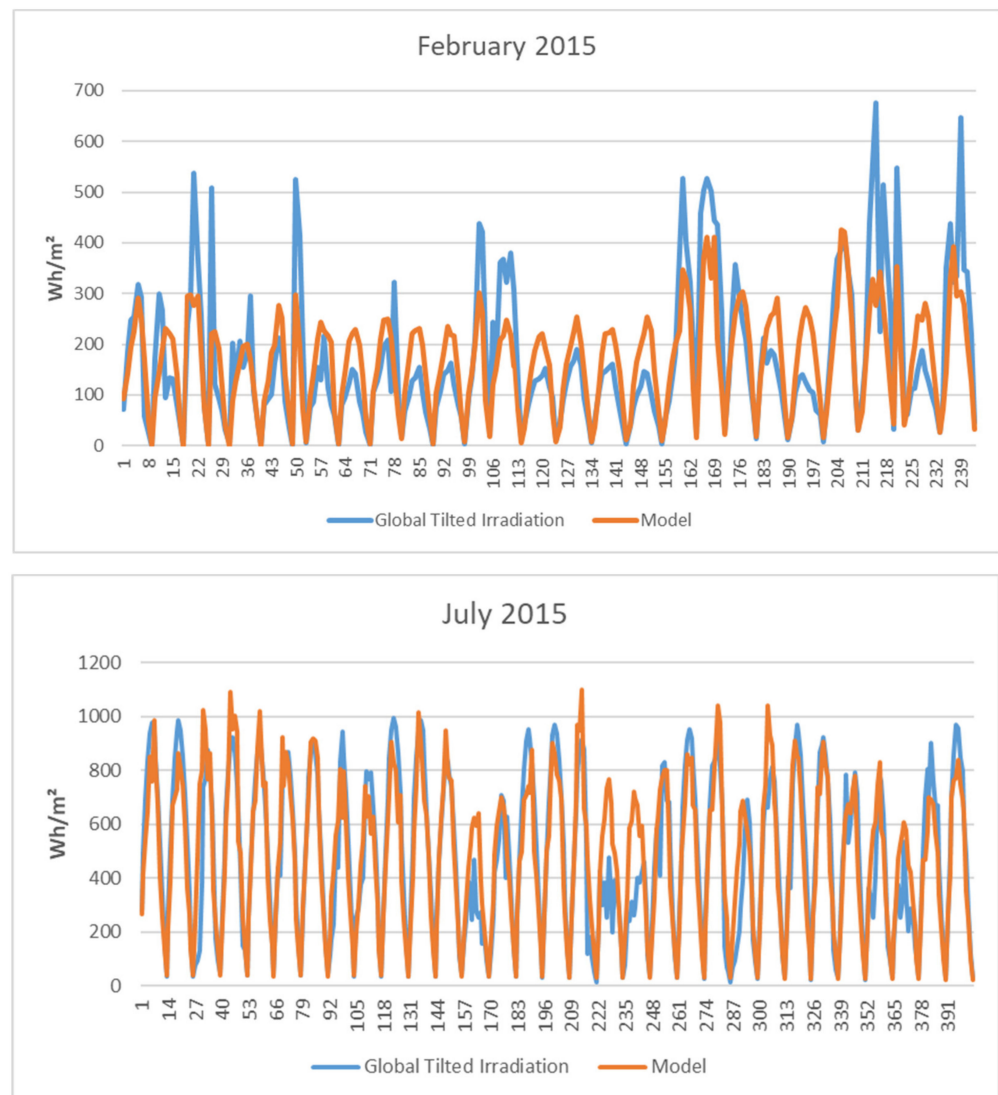


Figure 5. Real and simulated global tilted irradiation.

Differences in the summertime are comparably lower than differences in wintertime, which is significant for assessing our model. In the summertime, irradiation values are higher, so we have more resulting photovoltaic power. Smaller model errors mean that simulations reflect better the true conditions in these months; hence, we can consider them more reliable. Considering the February data, we see that we underestimate irradiation values on days with comparably intensive sunshine. This might be due to an underestimation of reflexive irradiation due to snow, for example.

5. Quality Control and Discussion

5.1. Quality Control

For evaluating the quality of our model, we have to rely on assessing the irradiation values due to the lack of accurate historic photovoltaic data. This case study's central and relevant findings can be summarized as follows: The annual errors' average is approximately zero. Hence, the model is unbiased, and the annual sum of irradiation from the model is roughly equal to the historical measurements. Even if the hourly estimates are incorrect, the overall simulated irradiation is correct, and any financial analysis will not be biased. However, in this section, we want to look at model errors to understand the model better and know how to interpret and treat the resulting simulations.

On a monthly level, relative differences between the modeled and true irradiation amounts are maximum around plus/minus 8%, as shown in Figure 6. Consequently, given unbiased temperature simulation, we slightly overestimate photovoltaic power amounts in March and underestimate them in August. The rest of the months, especially May and December, show comparably smaller relative differences.

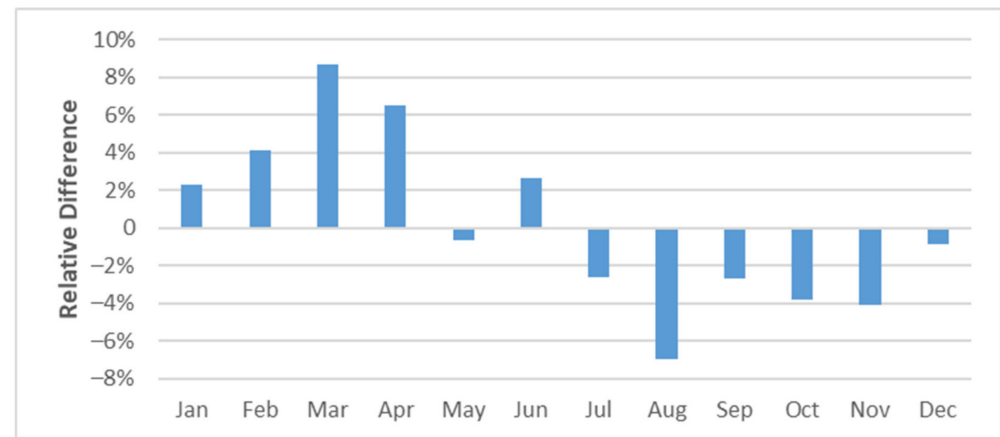


Figure 6. Relative difference between modeled and measured irradiation amounts (plain ground) on a monthly level for Ulm data between 2015 and 2018.

Eventually, we consider hourly data and compute (a) the relative standard errors (RSE), i.e., the differences' standard deviation divided by the respective hourly mean, as well as (b) the median absolute percentage error (MAPE). The second measure is computed by including values from the data set at a particular hour and month. For example, we analyze the error for Hour 5 in January. The same is done for the MAPE. We compute errors for irradiation on plain and on the tilted ground in order to see whether this has an effect on the error levels.

Given a data set of T observations, it is defined for E_t^{glo} as

$$MAPE\left(\left(\hat{E}_t^{glo}\right)_{t=1,\dots,T}, \left(E_t^{glo}\right)_{t=1,\dots,T}, hour, month\right) = \frac{1}{T} \sum_{j=1}^J \left| \frac{\hat{E}_j^{glo} - E_j^{glo}}{E_j^{glo}} \right|, \quad (12)$$

Thereby, \hat{E}_j^{glo} denotes the estimated irradiation value at time instance j , whereby only values at a chosen *hour* and in a specific *month* are considered. Consequently, J denotes the length of this subset. The MAPE for $E_t^{glo,tilt}$ is computed analogously. Thereby, as above, we assume a 30° southward tilt. We compute the median, not the mean, as the standard definition of the MAPE implies. When looking at the errors' distribution, we see a strong asymmetry as it is tilted to the right. Hence, the median is better to estimate the average absolute percentage error. Results for the average over hourly errors in monthly granularity for E_t^{glo} and $E_t^{glo,tilt}$ are given in Table 2.

The first thing we see is that errors indeed increase when considering $E^{glo,tilt}$ instead of E^{glo} . Additionally, we see that error measures are higher in the winter months than during summer. This is because irradiation values are higher during summer. Hence, the least squares optimization focuses more on those values than on the comparably lower values in winter. RSE values range between 25% and about 39% for E^{glo} and between 24% and 50% for $E^{glo,tilt}$, whereby the highest values are observed between October and February, where absolute values are lowest over the year. Average MAPE values for E^{glo} range between 17% and 34% with the same annual pattern, i.e., showing higher values during winter. MAPE values for $E^{glo,tilt}$ are slightly higher, ranging between 18% and 43%. Therefore, it is reasonable that both MAPE and RSE values are higher for $E^{glo,tilt}$ due to

the higher proportion of direct irradiation, which is more complex to model than diffuse irradiation as it is more stochastic.

Table 2. Errors of estimating global (tilted) irradiation in Ulm.

		E^{glo}					
		January	February	March	April	May	June
RSE		37.82%	38.94%	29.71%	29.17%	27.97%	25.11%
MAPE		29.01%	30.78%	25.96%	22.27%	19.85%	17.74%
		January	February	March	April	May	June
RSE		25.66%	27.50%	29.65%	34.96%	38.74%	32.06%
MAPE		17.18%	17.86%	21.55%	29.48%	32.05%	33.55%
		$E^{glo,tilt}$					
		January	February	March	April	May	June
RSE		49.98%	47.76%	37.57%	27.50%	29.15%	24.19%
MAPE		37.44%	39.14%	31.12%	24.62%	23.07%	18.95%
		January	February	March	April	May	June
RSE		25.37%	27.49%	31.86%	43.92%	48.80%	49.33%
MAPE		17.74%	20.01%	25.14%	37.65%	43.23%	39.13%

In order to understand the error levels summarized in Table 2, we have a closer look at the data. First, one has to consider that incidence angle, cloud coverage, temperature, and a parameter for the annual seasonality can only partly explain irradiation. In Figure 7, we plot irradiation versus cloud coverage. On the first day, the interdependencies are obvious as high cloudiness induces comparably low irradiation. On 5 and 6 February, global irradiation evolves constantly along with the daylight pattern, but cloudiness is oscillating significantly. However, when fitting Equation (10), the respective parameter α_2 is significant. One assumption is that hourly granularity already has too much aggregation and that temporary sunny periods within an hour overcompensate the cloudy phases. Additionally, global irradiation also considers indirect irradiation, which contributes to the total value also in clouded phases. Above that, there might be other influence factors or nonlinear interdependencies that have not been considered yet. However, the latter would mean the reliance on even more data, where the current data sets are sometimes hard to obtain. Working, for example, with 10-min data increases the amount of data by a factor 6. Moreover, data quality on an hourly level is comparably higher.

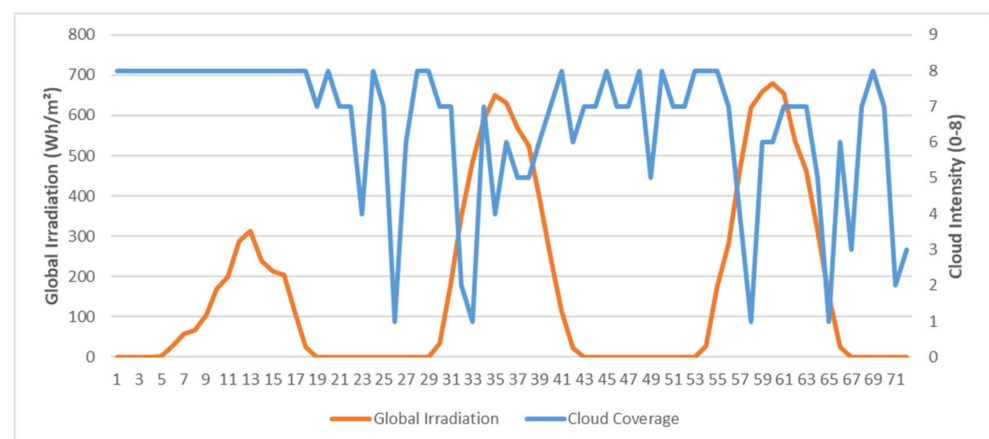


Figure 7. Global irradiation vs. cloudiness on a horizontal plane from 4 to 6 September 2014 in Ulm.

5.2. Discussion

For evaluating the potential of our model, we have to consider alternatives as well. Note that literature exclusively focuses on horizontal values, which is reasonable as tilt offers too much variety.

Politaki and Alouf's model [14] is thereby the model closest to our approach. Their model is based on the Clear Sky Model of Bird and Hulstrom [7], which is extended by a multiplicative stochastic factor to (a) include cloudiness and (b) allow for simulating irradiation values. However, it is hard to compare their model to ours regarding error numbers, as they show only root mean squared errors, which are even aggregated to an annual level. Given the strong seasonality of irradiation values, this information is only of limited value. Additionally, there is one significant drawback: as it is based on the Clear Sky Model, we require additional data such as pressure and aerosol depth. Especially the latter is hard to obtain—even more if historical observations are required. Lockart et al. [38] state that their simulation models achieve preciseness of 8%. However, they consider daily values and ignore intraday patterns, which are the challenging part here. Anand et al. [39] focus on the distributional properties. Trend and seasonal aspects are captured via a symmetric moving average given historical observations and then filtered. After that, first-order differences eliminate the last deterministic information. Simulation is purely done via the distribution fitted to the residuals of the previously described process. The advantage of their idea is that only historic irradiation values are required; however, limiting the annual and daily swing to a fixed average value narrows the range of possible scenarios significantly. Extraordinarily long/short summers or comparably cold/warm winters cannot be modeled in their setup. As we simulate temperature separately, our model accounts for different weather scenarios.

Benth et al. [9] use the sun intensity to identify and eliminate the seasonal component's influence. Afterward, they fit an autoregressive process, which can also be used for simulation purposes. Here, no local but regionally aggregated data on a daily level are used. Hence, this is also not an alternative to our approach.

These few examples illustrate that, despite a wide range of literature about this topic, it is hard to compare individual models. These vary in granularity (hourly, daily) and objective. However, we can state a few things. Compared to existing approaches which are close to our model [6,7], we can significantly reduce the amount of climate data required for calibration at the cost of reduced accuracy. One does not have to consider factors such as water vapor, ozone level, aerosol optical depth, and air pressure. These influence factors are all time-varying, and historical measurements (for model calibration purposes) are hard to obtain. Our model is much more user-friendly, as it is easier to handle and calibrate. Hourly error levels seem to be high at first sight. However, solar irradiation is a highly complex physical parameter. Moreover, our intention is not to model but to simulate data. Error values are also given to see whether numbers increase in the case of tilted irradiation. More importantly, the model is unbiased on an annual level, and model bias on a monthly level is relatively low.

6. Conclusions

Simulating photovoltaic power is required in the energy risk management of power utilities, power plant operators, and energy trading companies. Because of the relevance of this topic, literature already offers methods and tools to simulate photovoltaic power. This article adds a new parametric model to simulate solar irradiation and, consequently, photovoltaic power on an hourly level. Our approach is based on the idea that simulated global irradiation values are scaled values of extraterrestrial irradiation. The (hourly) scaling factors are obtained via linear functions that include factors such as temperature, cloudiness, or maximum possible irradiation. An annual trend is included as well. In order to derive corresponding solar power values, we integrate our approach into the cloudy sky model designed by Myers [6]. Our results are positive: the simulated irradiation pattern is similar to historical data. Additionally, we see no bias in the total solar power

output compared to the previous years. We have a clear and straightforward approach that includes the interdependencies between the individual weather factors. Hence, different possible weather scenarios are covered, but no complex physical knowledge or considerable computational power is needed. The application is therefore also suitable for optimizing the operation of isolated photovoltaic grids.

However, from a modeling point of view, there is still some research to be done in the future. Relative errors of solar irradiation models are relatively high compared to other fields of application. Especially in the wintertime, our model lacks some precision. Future research needs to verify whether nonlinear models should be considered as well or whether other weather factors such as precipitation and aerosol density should be included.

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Appendix A

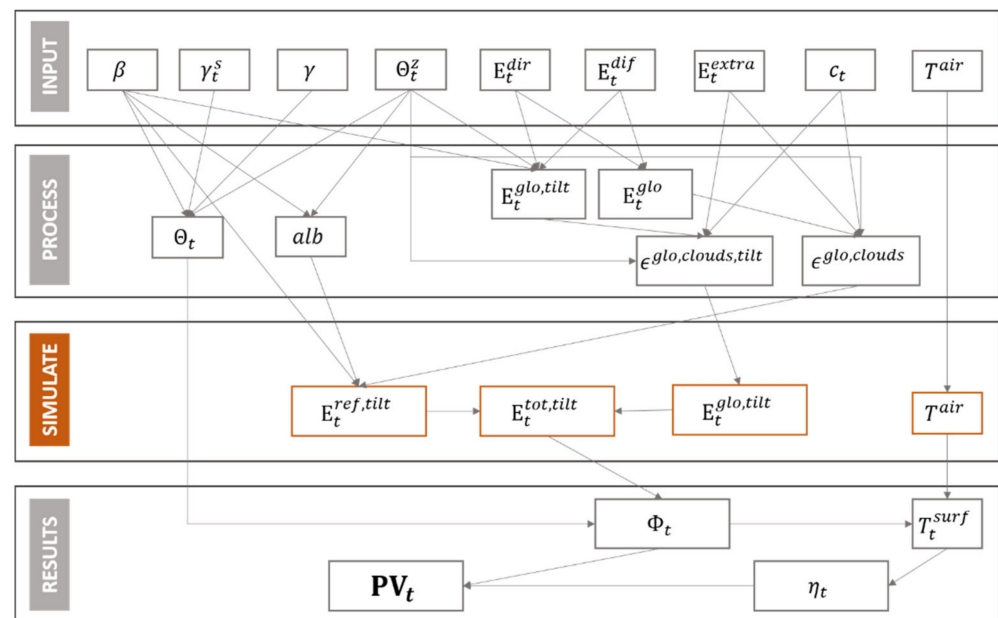


Figure A1. Parameter interactions.

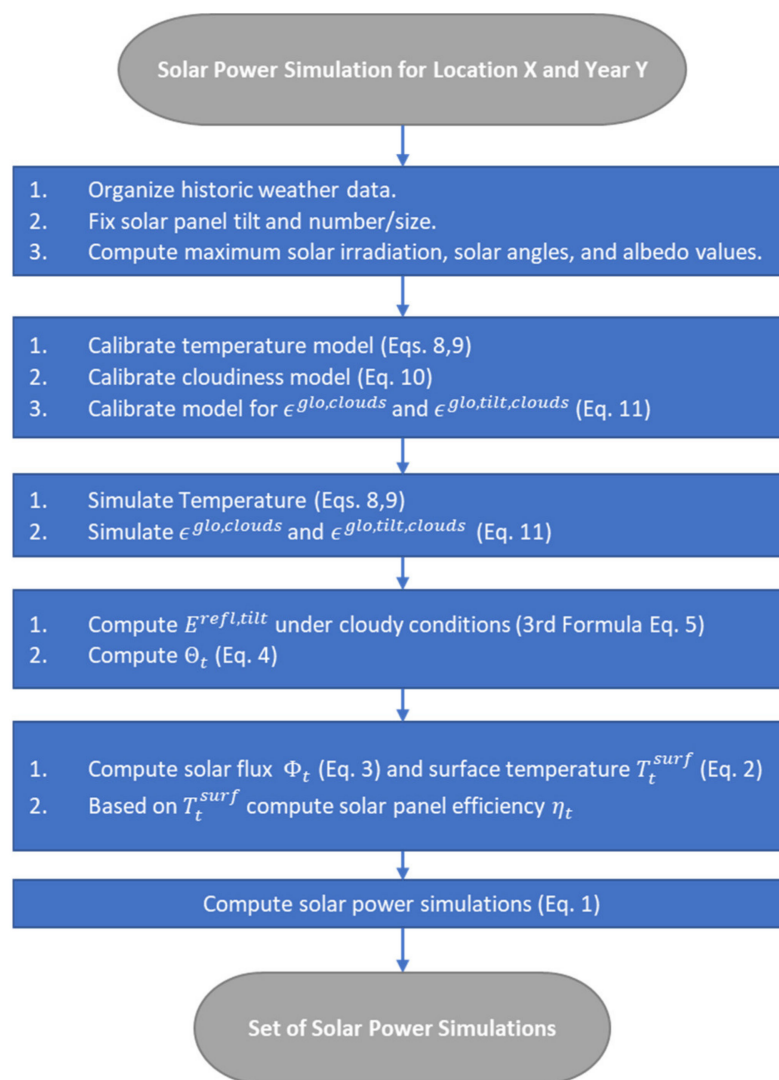


Figure A2. Flow chart of the proposed solar irradiation model.

Table A1. Cloud coverage transition matrix for Ulm, Germany.

From/To	0	1	2	3	4	5	6	7	8
0	68.93%	7.95%	4.70%	3.41%	2.78%	2.58%	2.45%	4.26%	2.94%
1	31.02%	12.98%	9.95%	7.90%	8.29%	7.51%	4.98%	9.27%	8.10%
2	18.11%	11.19%	12.31%	10.99%	8.95%	10.27%	7.53%	11.50%	9.16%
3	12.40%	7.66%	9.66%	9.66%	12.31%	11.58%	9.39%	16.13%	11.21%
4	8.62%	6.81%	7.80%	10.92%	12.56%	13.63%	12.64%	15.68%	11.33%
5	6.38%	4.77%	7.18%	9.09%	11.22%	12.46%	12.10%	20.09%	16.72%
6	4.36%	3.37%	6.20%	7.66%	9.64%	10.03%	13.14%	26.67%	18.94%
7	1.99%	1.82%	1.60%	2.73%	3.08%	4.14%	6.34%	32.27%	46.03%
8	0.58%	0.46%	0.47%	0.65%	0.80%	1.06%	1.38%	14.44%	80.16%

Table A2. Parameters for estimating E^{glo} of Ulm.

Hour	Intercept	c_t	$90^\circ - \Theta_z$	E^{extra}	T^{air}	Seasonality
6	0.8664	−0.0447	−0.1085	0.0041	0.0052	0.0565
7	0.9336	−0.0466	−0.1556	0.0061	0.0038	0.1519
8	1.1537	−0.0460	−0.0251	0.0002	0.0044	0.2764
9	1.1165	−0.0449	−0.0077	−0.0003	0.0061	0.2492
10	0.8488	−0.0423	−0.0356	0.0014	0.0060	0.1797
11	0.7877	−0.0377	−0.0230	0.0009	0.0083	0.1552
12	0.7460	−0.0325	−0.0197	0.0008	0.0082	0.1694
13	0.7132	−0.0306	−0.0154	0.0006	0.0078	0.1571
14	0.6043	−0.0285	−0.0189	0.0008	0.0071	0.1300
15	0.5085	−0.0232	−0.0219	0.0009	0.0062	0.1297
16	0.2547	−0.0184	−0.0654	0.0031	0.0042	0.1147
17	0.0964	−0.0134	−0.0141	0.0010	0.0020	0.0379
18	0.0256	−0.0084	−0.0180	0.0013	0.0008	0.0592
19	−0.0276	−0.0044	−0.0201	0.0015	−0.0005	0.0320

Table A3. Parameters for estimating $E^{glo, tilt}$ of Ulm.

Hour	Intercept	c_t	$90^\circ - \Theta_z$	E^{extra}	T^{air}	Seasonality
6	0.3930	0.0063	−0.2051	0.0089	−0.0011	−0.1730
7	0.8892	−0.0336	−0.0559	0.0017	0.0007	0.0954
8	0.9814	−0.0427	−0.0437	0.0013	0.0044	0.1819
9	1.0066	−0.0452	−0.0423	0.0014	0.0060	0.2181
10	0.7323	−0.0451	−0.0583	0.0027	0.0060	0.1548
11	0.7920	−0.0403	−0.0414	0.0018	0.0117	0.1808
12	0.9114	−0.0354	−0.0332	0.0012	0.0130	0.2369
13	0.9308	−0.0398	−0.0267	0.0009	0.0119	0.1983
14	0.8582	−0.0449	−0.0334	0.0013	0.0106	0.1491
15	0.6622	−0.0401	−0.0485	0.0021	0.0078	0.1243
16	0.2767	−0.0271	−0.1058	0.0050	0.0045	0.0870
17	0.1091	−0.0173	−0.1230	0.0058	0.0018	0.0498
18	0.01606	−0.0060	−0.0104	0.0009	0.0005	0.0273
19	−0.0280	0.0007	−0.0097	0.0009	−0.0007	0.0174

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